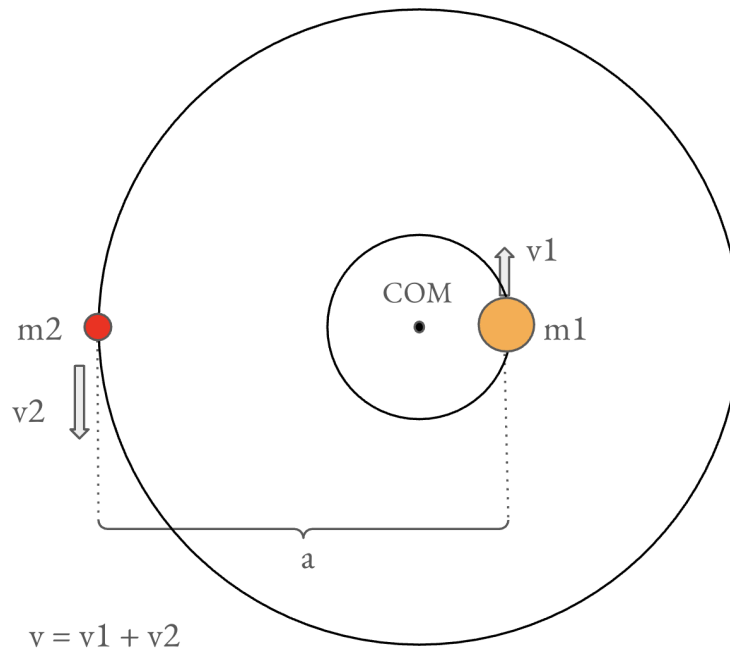
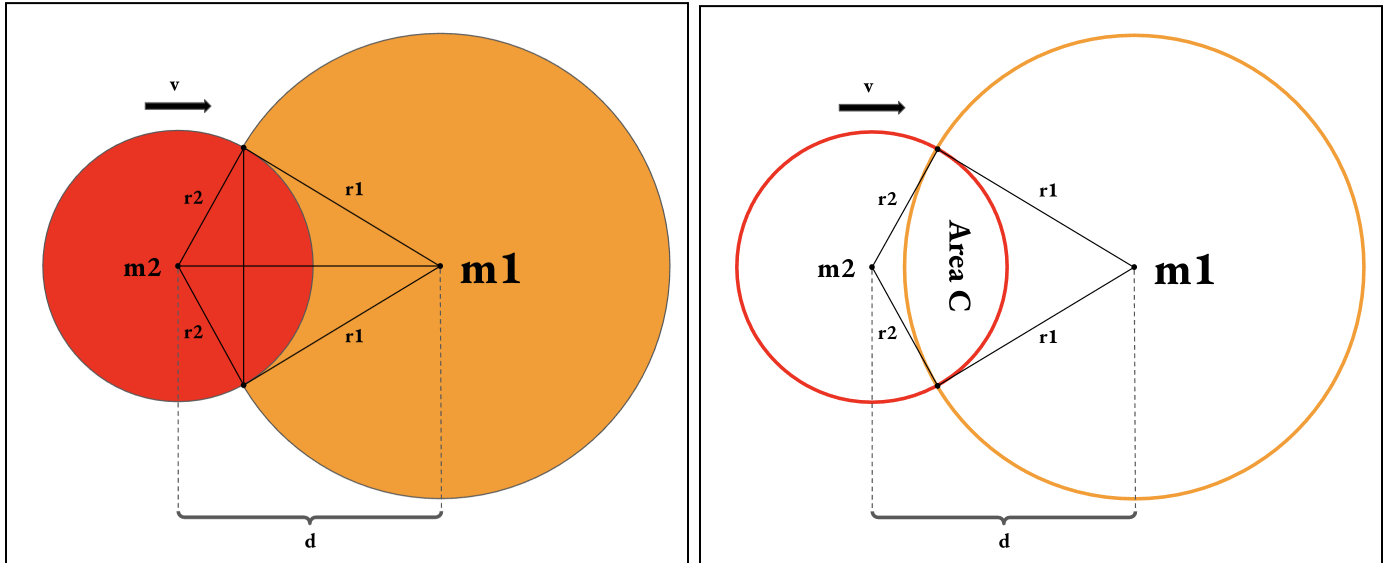


(1) Major Assumptions:

- Uniform Luminosity Distribution (No limb darkening, luminosity/projected area = const)
- Circular Orbit ($e = 0$)
- Orbital Inclination is 90 deg
- Distance to System $\gg$ Orbital Radii
- Celestial bodies m_1 and m_2 is perfectly spherical with projected areas being perfect circles



$$v_{rel} = \sqrt{\frac{G(m_1+m_2)}{a}} = \left| \vec{v}_1 \right| + \left| \vec{v}_2 \right|$$

$$v_1 = \frac{m_2}{m_1+m_2} \sqrt{\frac{G(m_1+m_2)}{a}}, \quad v_2 = \frac{m_1}{m_1+m_2} \sqrt{\frac{G(m_1+m_2)}{a}}$$

$$Tangential\ Velocity = v(t)$$

$$v(t) = v_{rel} \cos(\omega t - \phi), \quad \omega = \frac{2\pi}{P}, \quad \phi = \arcsin\left(\frac{r_1+r_2}{a}\right)$$

$$\text{Eclipse begins at } t = 0$$

$$x(t) = \int v(t) dt = \frac{v_{rel}}{\omega} \sin(\omega t - \phi)$$

$$\Rightarrow x(t) = \frac{v_{rel}}{\omega} \sin(\omega t - \phi) + (r_1 + r_2)$$

$$d(t) = \left| (r_1 + r_2) - x(t) \right| = \left| -\frac{v_{rel}}{\omega} \sin(\omega t - \phi) \right|$$

$$2r_2 = \frac{v_{rel}}{\omega} \sin(\omega t - \phi) + (r_1 + r_2)$$

$$r_2 - r_1 = \frac{v_{rel}}{\omega} \sin(\omega t - \phi)$$

$$\frac{\arcsin\left((r_2-r_1)\frac{\omega}{v_{rel}}\right)+\phi}{\omega} = t_1$$

$$2r_2 + (2r_1 - 2r_2) = \frac{v_{rel}}{\omega} \sin(\omega t - \phi) + (r_1 + r_2)$$

$$r_1 - r_2 = \frac{v_{rel}}{\omega} \sin(\omega t - \phi)$$

$$\frac{\arcsin\left((r_1-r_2)\frac{\omega}{v_{rel}}\right)+\phi}{\omega} = t_2$$

$$2r_1 + 2r_2 = -\frac{v_{rel}}{\omega} \sin(\omega t - \phi) + (r_1 + r_2)$$

$$\frac{\arcsin\left((r_1+r_2)\frac{\omega}{v_{rel}}\right)+\phi}{\omega} = t_3$$